

PowerPoint Lectures
to accompany
Physical Science, 8e

Chapter 8
Atoms and Periodic
Properties

*** Core Concept**

Different fields of study contributed to the development of a model of the atom.

- *Dalton - atoms indivisible
- *Thomson and Millikan experiments
 - *Electron mass very small, no measurable volume
 - *What is the nature of an atom's positive charge?
- *Thomson's "Plum pudding" model
 - *Electrons embedded in blob of positively charged matter like "raisins in plum pudding"

*** Early Models of the Atom**

*The Nucleus

- Ernest Rutherford (1907)
- * Scattered alpha particles off gold foil
- * Most passed through without significant deflection
- * A few scattered at large angles
- * Conclusion: an atom's positive charge resides in a small, massive nucleus
- * Later: positive charges = protons
- * James Chadwick (1932): also neutral neutrons in the nucleus

*The Nuclear Atom

- * Atomic number
 - * Number of protons in nucleus
 - * Elements distinguished by atomic number
 - * 113 elements identified
 - * Number of protons = number of electrons *in neutral atoms*
- * Isotopes
 - * Same number of protons; different number of neutrons

*Atomic Symbols and Masses

- Mass number
 - * Number of protons + neutrons
- Atomic mass units (u)
 - * 1/12 of carbon-12 isotope mass
- Atomic weight
 - * Atomic mass of an element, averaged over naturally occurring isotopes

Predictions of classical theory
*Electrons orbit the nucleus
*Curved path = acceleration
*Accelerated charges radiate
*Electrons lose energy and spiral into nucleus
*Atoms cannot exist!
Experiment - atoms do exist
⇒ New theory needed

*Classical "Atoms"

*The Quantum Concept

*Max Planck (1900)
*Introduced quantized energy
*Einstein (1905)
*Light made up of quantized photons
*Higher frequency photons = more energetic photons

*Atomic Spectra

Blackbody radiation
*Continuous radiation distribution
*Depends on temperature of radiating object
*Characteristic of solids, liquids and dense gases
Line spectrum
*Emission at characteristic frequencies
*Diffuse matter: incandescent gases
*Illustration: Balmer series of hydrogen lines

*Bohr's Theory

Three rules:

1. Electrons only exist in certain allowed orbits
2. Within an orbit, the electron does not radiate
3. Radiation is emitted or absorbed when changing orbits

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Chapter 13
Nuclear Reactions

Nuclear reactions involve changes in the nucleus of the atom.

*Core Concept

* Nuclear Radioactivity

- * Natural radioactivity
 - * Spontaneous emission of particles or energy from an unstable nucleus
 - * Discovered by Becquerel
- * Three types of radioactive decay
 - * Alpha decay (He-nucleus)
 - * Beta decay (high energy electron)
 - * Gamma decay (high energy electromagnetic radiation)

* Nuclear Equations

- * Atomic number = number of protons in nucleus
- * Isotopes; same atomic number; different number of neutrons
- * Mass number = number of nucleons (protons and neutrons) in nucleus
- * Nuclear reactions
 - * Represented by balanced equations
 - * Charge conserved
 - * Mass number conserved

* The Nature of the Nucleus

- * Strong nuclear force
 - * Binds protons and neutrons
 - * Very short ranged, less than 10^{-15} m
 - * Overcomes proton-proton Coulomb repulsion
- * Nuclear shell model
 - * Nucleon quantum energy levels
 - * Maximum stability for nucleon number = 2, 8, 20, 28, 50, 82 or 126
- * Band of stability

1. Atomic number > 83: unstable
2. Nucleon number = 2, 8, 20, 28, 50, 82 or 126: added stability
3. Pairs of protons and pairs of neutrons: added stability
 - Odd number of both protons and neutrons less stable
4. Neutron: proton ratios for added stability
 - 1:1 in isotopes with up to 20 protons
 - 1+increasing:1 with increasingly heavy isotopes

*Generalizations - Nuclear Stability

*Types of Radioactive Decay

1. Alpha emission
 - Expulsion of helium nucleus
 - Least penetrating: stopped by paper
$${}_{86}^{222}\text{Rn} \rightarrow {}_{84}^{218}\text{Po} + {}_2^4\text{He}$$
2. Beta emission
 - Expulsion of an electron
 - More penetrating: 1 cm of aluminum
$${}_{6}^{14}\text{C} \rightarrow {}_{7}^{14}\text{N} + {}_{-1}^0\text{e}$$
3. Gamma decay
 - Emission of a high energy photon
 - Most penetrating: 5 cm of lead
$${}_{86}^{222}\text{Rn}^* \rightarrow {}_{86}^{222}\text{Rn} + {}_0^0\gamma$$

*Radioactive Decay Series

- *One radioactive nucleus decays to a 2nd, which decays to a 3rd, which...
- *Three naturally occurring series
 - *Thorium-232 to lead-208
 - *Uranium-235 to lead-207
 - *Uranium-238 to lead-206

* Half-life

- * Time required for 1/2 of a radioactive sample to decay
- * Example: 1 kg of an unstable isotope with a one-day half-life
 - * After 1 day: 500 g remain
 - * After 2 days: 250 g remain
 - * After 3 days: 125 g remain
- * U-238 decay series: wide half-life variation

* Measurement of Radiation

- Measurement methods
- * Ionization counters
 - * Detect ions produced by radiation
 - * Example: Geiger counter
 - * Scintillation counters
 - * Rely on flashes of light produced as radiation strikes a phosphor
 - * Zinc sulfide: phosphor used in TV picture tubes

* Radiation Units

- * Measured at the source
 - * Activity: number of disintegrations per unit time
 - * Units: Becquerel (SI unit), Curie, ...
- * Measured where absorbed
 - * Human exposure: rem
 - * SI unit: millisievert
 - * rad: radiation absorbed dose (unit = gray)
 - * Dosage related to effects on organism

* Radiation Exposure

- * Natural radioactivity
 - * 100-500 mrem/yr
- * Sources
 - * Cosmic rays from outer space
 - * Earth's residual radioactivity
 - * Medical x-rays, TVs, ...
- * Consequences
 - * DNA disruption
 - * Free radical production
- * Threshold versus linear exposure models

* Nuclear Energy

- * Interconversion of mass and energy
- * Mass defect
 - * Difference between masses of reactants and products
- * Binding energy
 - * Energy required to break a nucleus into individual protons and neutrons
- * Ratio: binding energy to nucleon number
 - * Iron-56 = most stable nucleus

* Nuclear Fission

- * Heavy nuclei splitting into lighter ones
- * Chain reactions
 - * Possible when one reaction can lead to others
 - * One neutron in, two or more out
- * Critical mass
 - * Sufficient mass and concentration to produce a chain reaction

* Nuclear Power Plants

- * Rely on controlled fission chain reactions
- * Steel vessel contains fuel rods and control rods
- * Full plant very intricate
 - * Containment and auxiliary buildings necessary
- * Spent fuel rods
 - * Contain fissionable materials U-235, Pu-239
 - * Disposal issues not settled

* Nuclear Fusion

- * Less massive nuclei forming more massive nuclei
- * Energy source for Sun and other stars
- * Requirements for fusion
 - * High temperature
 - * High density
 - * Sufficient confinement time
- * Controlled fusion
 - * Magnetic confinement
 - * Inertial confinement

* Ultimately connected to origins of the Universe and the life cycles of stars

* Big Bang theory

- * Incredibly hot, dense primordial plasma cools, creating protons and neutrons
- * Continued cooling leads to hydrogen atoms which collapse gravitationally into 1st generation stars

* Stellar evolution

- * Interior temperatures and densities suitable for fusion of heavy elements beyond hydrogen and helium
- * Certain massive stars explode in supernovae, spreading heavy elements (some radioactive)
- * Ultimate source: gravitational attraction!

* Source of Nuclear Energy
